# Markscheme 

May 2018

Physics

Higher level

Paper 3

This markscheme is the property of the International
Baccalaureate and must not be reproduced or distributed to any other person without the authorization of the IB Global Centre, Cardiff.

Section A

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 1. | a |  | smooth line, not kinked, passing through all the error bars $\checkmark$ |  | 1 |
| 1. | b | i | $0.84 \pm 0.03$ «s》 ${ }^{\text {s }}$ | Accept any value from the range: 0.81 to 0.87 . <br> Accept uncertainty 0.03 OR 0.025. | 1 |
| 1. | b | ii | $\begin{aligned} & K=\sqrt{0.005} \times 0.84=0.059 \\ & \text { « } \frac{\Delta K}{K}=\frac{\Delta P}{P} » \\ & \Delta K=\frac{0.03}{0.84} \times 0.0594=0.002 \\ & « K=(0.059 \pm 0.002) » \end{aligned}$ <br> uncertainty given to 1 sf $\checkmark$ | Allow ECF [3 max] if 10T is used. Award [3] for BCA. | 3 |
| 1. | b | iii | $s T^{\frac{1}{2}},$ | Accept $s \sqrt{T}$ or in words. | 1 |
| 1. | c |  | straight $\boldsymbol{A N D}$ ascending line $\checkmark$ through origin $\checkmark$ |  | 2 |
| 1. | d |  | $K=\sqrt{\text { slope }} \checkmark$ |  | 1 |


| 2. | a |  |  |  |
| :--- | :--- | :--- | :--- | :--- | :--- |

## Section B

## Option A — Relativity

| 3. | $\mathbf{a}$ |  | magnetic field $\checkmark$ |  |  |
| :--- | :--- | :--- | :--- | :--- | :--- |
| 3. | $\mathbf{b}$ | $\mathbf{i}$ | «according to $Y$ » the positive charges are moving «to the right» $\checkmark$ <br> $d$ decreases $\checkmark$ | For MP1, movement of positive charges <br> must be mentioned explicitly. |  |
| 3. | $\mathbf{b}$ | $\mathbf{i i}$ | positive charges are moving, so there is a magnetic field $\checkmark$ <br> the density of positive charges is higher than that of negative charges, <br> so there is an electric field $\checkmark$ | The reason must be given for each <br> point to be awarded. | $\mathbf{2}$ |


| 4. | a | i | $\text { " } \frac{10^{4}}{0.995 \times 3 \times 10^{8}}=>34<\mu s » \checkmark$ | Do not accept $10^{4} / \mathrm{c}=33 \mu \mathrm{~s}$. | 1 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 4. | a | ii | time is much longer than 10 times the average life time «so only a small proportion would not decay» $\checkmark$ |  | 1 |
| 4. | b | i | $\begin{aligned} & \gamma=10 \checkmark \\ & \Delta t_{0}=« \frac{\Delta t}{\gamma}=\frac{34}{10}=» 3.4<\mu s » \end{aligned}$ |  | 2 |
| 4. | b | ii | the value found in (b)(i) is of similar magnitude to average life time $\checkmark$ significant number of muons are observed on the ground $\checkmark$ «therefore this supports the special theory" |  | 2 |


| 5. | a |  | straight line with negative gradient with vertical intercept at $c t=1.2$ «km» $\checkmark$ through $(-0.6,2.2)$ ie gradient $=-1.67 \checkmark$ | Tolerance: Allow gradient from interval -2.0 to -1.4, (at ct $=2.2, x$ from interval 0.5 to 0.7). <br> If line has positive gradient from interval 1.4 to 2.0 and intercepts at $c t=1.2 \mathrm{~km}$ then allow [1 max]. | 2 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 5. | b |  | line for the flash of light from A correctly drawn $\checkmark$ <br> line for the flash of light of B correctly drawn $\checkmark$ <br> correct reading taken for time of intersection of flash of light and path of B, $c t=2.4$ «km» $\checkmark$ | Accept values in the range: 2.2 to 2.6. | 3 |

(continued...)
(Question 5 continued)

| 5. | b |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 5. | C |  | the two events take place in the same point in space at the same time $\checkmark$ so all observers will observe the two events to be simultaneous / so zero difference $\checkmark$ | Award the second MP only if the first MP is awarded. | 2 |
| 5. | d |  | $\begin{aligned} & u^{\prime}=\frac{-0.6-0.8}{1-(-0.6) \times 0.8} \\ & =«-» 0.95 « c » \checkmark \end{aligned}$ |  | 2 |


| 6. | a | i | «-» 29.8 «MeVc-1» ${ }^{-1}$ |  | 1 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 6. | a | ii | $\begin{aligned} & E_{\pi}=\sqrt{p_{\mu}^{2} c^{2}+m_{\mu}^{2} c^{4}}+p_{\vee} c \text { OR } E_{\mu}=109.8 « \mathrm{MeV} » \\ & E_{\pi}=« \sqrt{29.8^{2}+105.7^{2}}+29.8=» 139.6 « \mathrm{MeV} » \end{aligned}$ | Final value to at least 3 sig figs required for mark. | 2 |
| 6. | b |  | $139.6 \mathrm{MeVc}^{-2} \checkmark$ | Units required. Accept $140 \mathrm{MeVc}^{-2}$. | 1 |


| 7. | a | $\Delta f \propto f \checkmark$ <br> therefore the change is $<->3 \Delta f \checkmark$ |  | 2 |
| :---: | :---: | :---: | :---: | :---: |
| 7. | b | $\begin{aligned} & g=« c^{2} \frac{\Delta f}{f \Delta h}=»\left(3 \times 10^{8}\right)^{2} \frac{170}{5.0 \times 10^{14} \times 10000} \\ & g=3.1<\mathrm{ms}^{-2} » \end{aligned}$ | If POT mistake, award [0]. Award [2] for BCA. | 2 |
| 7. | c | the mass of the planet warps spacetime around itself $\checkmark$ the light will follow the shortest path in spacetime «which is curved» $\checkmark$ |  | 2 |

## Option B - Engineering physics

| 8. | a |  | $\begin{aligned} & \Gamma «=F r=50 \times 2 »=100 « \mathrm{Nm} » \\ & \alpha «=\frac{\Gamma}{I}=\frac{100}{450} »=0.22 « \mathrm{rads}^{-2} » \end{aligned}$ | Final value to at least 2 sig figs, $\mathbf{O R}$ clear working with substitution required for mark. | 2 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 8. | b | i | $\begin{aligned} & « \omega_{t}^{2}-\omega_{0}^{2}=2 \alpha \Delta \theta » \\ & « \omega_{t}^{2}-0=2 \times 0.22 \times 2 \pi » \\ & \omega_{t}=1.7 \text { «rads }^{-1} » \checkmark \end{aligned}$ | Accept BCA, values in the range: 1.57 to 1.70 . | 1 |
| 8. | b | ii | $\begin{aligned} & « L=I \omega=450 \times 1.66 » \\ & =750<\mathrm{kgm}^{2} \mathrm{rads}^{-1} » \end{aligned}$ | Accept BCA, values in the range: 710 to 780. | 1 |
| 8. | c |  | $\begin{aligned} & « I=450+m r^{2} » \\ & I «=450+30 \times 2^{2} »=570 « \mathrm{kgm}^{2} » \checkmark \\ & « L=570 \times \omega=747 » \\ & \omega=1.3 « \mathrm{rads}^{-1} » \checkmark \end{aligned}$ | Watch for ECF from (a) and (b). <br> Accept BCA, values in the range: 1.25 to 1.35 . | 2 |

(Question 8 continued)

| 8. | $\mathbf{d}$ | $\mathbf{i}$ | moment of inertia will decrease $\checkmark$ <br> angular momentum will be constant «as the system is isolated» $\checkmark$ <br> «so the angular speed will increase» |  |
| :--- | :--- | :--- | :--- | :--- |
| 8. | $\mathbf{d}$ | $\mathbf{i i}$ | $\omega_{t}=1.66$ from bi $A N D W=\Delta E_{k} \checkmark$ <br> $W=\frac{1}{2} \times 450 \times 1.66^{2}-\frac{1}{2} \times 570 \times 1.31^{2}=131 « J » \checkmark$ | ECF from 8(b)(i). |


| 9. | a | $\begin{aligned} & « p_{1} V_{1}^{\frac{5}{3}}=p_{2} V_{2}^{\frac{5}{3}} » \\ & 1.1 \times 10^{5} \times 5^{\frac{5}{3}}=p_{2} \times 2^{\frac{5}{3}} \checkmark \\ & p_{2} «=\frac{1.1 \times 10^{5} \times 5^{\frac{5}{3}}}{2.5^{\frac{5}{3}}} \geqslant=5.066 \times 10^{5} \text { «Pa } » \end{aligned}$ | Volume may be in litres or $\mathrm{m}^{3}$. <br> Value to at least 2 sig figs, OR clear working with substitution required for mark. | 2 |
| :---: | :---: | :---: | :---: | :---: |

(continued...)
(Question 9 continued)


| 10. | a |  | in laminar flow, the velocity of the fluid is constant «at any point in the fluid» «whereas it is not constant for turbulent flow" $\downarrow$ | Accept any similarly correct answers. | 1 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 10. | b |  | $P_{S}=P_{T}$ «as both are exposed to atmospheric pressure» $\checkmark$ then $V_{T}=0$ «if the surface area of the reservoir is large» $\checkmark$ $\begin{aligned} & « \frac{1}{2} \rho v_{S}^{2}+\rho g z_{S}=\rho g z_{T} " \\ & \frac{1}{2} v_{S}^{2}=g\left(z_{T}-z_{S}\right)=g H \end{aligned}$ <br> and so $v_{s}=\sqrt{2 g H}$ | MP1 and MP2 may be implied by the correct substitution showing line 3 in the mark scheme. <br> Do not accept simple use of $v=\sqrt{2 \mathrm{as}}$. | 3 |
| 10. | c | i | $R=\frac{59.4 \times 0.6 \times 1 \times 10^{3}}{1.31 \times 10^{-3}}=2.72 \times 10^{7}$ | Accept use of radius 0.3 m giving value $1.36 \times 10^{7}$. | 1 |
| 10. | c | ii | as $R>1000$ it is not reasonable to assume laminar flow $\checkmark$ |  | 1 |
| 11. | a |  | damped oscillation / OWTTE |  | 1 |
| 11. | b | i | $E$ «= $\frac{1}{2} \times 30 \times \pi^{2} \times 0.8^{2} »=95 《 ل$ » $\checkmark$ | Allow initial amplitude between 0.77 to 0.80 , giving range between: 88 to 95 J . | 1 |
| 11. | b | ii | $\begin{aligned} & \Delta E=95-\frac{1}{2} \times 30 \times \pi^{2} \times 0.72^{2}=18 « \mathrm{~J}> \\ & Q=« 2 \pi \frac{95}{18}=>33 \end{aligned}$ | Accept values between 0.70 and 0.73 , giving a range of $\Delta E$ between 22 and 9 , giving $Q$ between 27 and 61 . <br> Watch for ECF from (b)(i). | 2 |

## Option C - Imaging

| 12. | a | i | image is real «as projected on a screen» $\checkmark$ |  | 1 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 12. | a | ii | $\begin{aligned} & «-\frac{18}{u}=-0.40 » \\ & u=45 \\ & \frac{1}{45}+\frac{1}{18}=\frac{1}{f} \end{aligned}$ <br> OR $\begin{aligned} & f=13 « \mathrm{~cm} » \\ & P=\frac{1}{f}=« \frac{1}{13} »=0.078<\mathrm{cm}^{-1} » \end{aligned}$ | Accept answer 7.7«D». | 3 |
| 12. | a | iii | refractive index depends on wavelength $\checkmark$ <br> light of different wavelengths have different focal points / refract differently $\checkmark$ there will be coloured fringes around the image / image will be blurred $\checkmark$ |  | 3 |
| 12. | b |  | any 2 correct rays to find image from lens $1 \checkmark$ ray to locate $F_{2} \checkmark$ <br> Focal length $=«-» 70 « \mathrm{~cm} » \downarrow$ | Accept values in the range: 65 cm to 75 cm . <br> Accept correct MP3 from accepted range also if working is incorrect or unclear, award [1]. | 3 |

(continued...)
(Question 12 continued)


| 13. | a |  | $\begin{gathered} « \sin c=\frac{1.34}{1.56} \text { " } \\ c=59.2 \ll \gg \end{gathered}$ | Accept values in the range: 59.0 to 59.5. <br> Accept answer 1.0 rad. | 1 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 13. | b |  | optic fibres are not susceptible to earthing problems $\checkmark$ optic fibres are very thin and so do not require the physical space of electrical cables $\checkmark$ <br> optic fibres offer greater security as the lines can not be tapped $\checkmark$ optic fibres are not affected by external electric/magnetic fields/interference $\checkmark$ optic fibres have lower attenuation than electrical conductors / require less energy $\checkmark$ <br> the bandwidth of an optic fibre is large and so it can carry many communications at once/in a shorter time interval /faster data transfer $\checkmark$ |  | 2 max |
| 13. | c | i | a signal that is wider and lower, not necessarily rectangular, but not a larger area $\checkmark$ |  | 1 |
| 13. | c | ii | $\begin{aligned} & \text { attenuation }=-1.24 \times 3.4 \text { « }=-4.216 \mathrm{~dB} » \\ & -4.216=10 \log \frac{I}{15} \checkmark \\ & I=5.68 \text { «mW» } \end{aligned}$ | Need negative attenuation for MP1, may be shown in MP2. <br> For MP3 answer must be less than 15 mW (even with ECF) to earn mark. Allow [3] for BCA. | 3 |

(continued...)
(Question 13 continued)


| 14. | a |  | crystal vibration /piezo-electric effect $\checkmark$ caused by an alternating potential difference is applied across a crystal $\checkmark$ |  | 2 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 14. | b |  | ADVANTAGES <br> the wavelength must be less than the size of the object being imaged to avoid diffraction effects $\checkmark$ <br> the frequency must be high to ensure several full wavelengths in the pulse $\checkmark$ <br> DISADVANTAGES <br> the depth of the organ being imaged must be considered (no more than 200 wavelengths) $\checkmark$ <br> attenuation increases at higher frequencies $\checkmark$ | [1] for advantages, [1] for disadvantages. | 2 max |
| 14. | C |  | X-rays are an ionizing radiation and so might cause harm to the developing fetus. OR <br> there are no known harmful effects when using ultrasound $\checkmark$ | Ignore "moving images by ultrasound". | 1 |

(continued...)
(Question 14 continued)

| 14. | d | i | $\rho=\frac{1.99 \times 10^{6}}{1.73 \times 10^{3}}=1.15 \times 10^{3}<\mathrm{kgm}^{-3} » \checkmark$ |  | 1 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 14. | d | ii | $\begin{aligned} & F=\frac{\left(1.99 \times 10^{6}-4.3 \times 10^{2}\right)^{2}}{\left(1.99 \times 10^{6}+4.3 \times 10^{2}\right)^{2}}=1.0 \\ & F=\frac{\left(1.48 \times 10^{6}-1.99 \times 10^{6}\right)^{2}}{\left(1.48 \times 10^{6}+1.99 \times 10^{6}\right)^{2}}=0.02 \end{aligned}$ <br> almost $100 \%$ of the ultrasound will be reflected from the air-skin surface $\mathbf{O R}$ almost none is transmitted $\checkmark$ <br> whereas only $2 \%$ will be reflected from the gel-skin surface and so a much greater proportion is transmitted $\checkmark$ | Need to explain that more is transmitted through gel-skin surface for MP4. | 4 |

## Option D - Astrophysics

| 15. | $\mathbf{a}$ | $\mathbf{i}$ | a galaxy is much larger in size than a solar system $\checkmark$ <br> a galaxy contains more than one star system / solar system $\checkmark$ <br> a galaxy is more luminous $\checkmark$ | Any other valid statement. |
| :--- | :--- | :--- | :--- | :--- | :--- |
| $\mathbf{1 5 .}$ | $\mathbf{a}$ | $\mathbf{i i}$ | a comet is a small icy body whereas a planet is mostly made of rock or gas $\checkmark$ <br> a comet is often accompanied by a tail/coma whereas a planet is not $\checkmark$ <br> comets (generally) have larger orbits than planets $\checkmark$ <br> a planet must have cleared other objects out of the way in its orbital neighbourhood $\checkmark$ | $\mathbf{1 ~ m a x ~}$ |


| 16. | $\mathbf{a}$ | $\mathbf{i}$ | the wavelengths of the dips correspond to the wavelength in the <br> emission spectrum $\checkmark$ <br> the absorption lines in the spectrum of star X suggest it contains predominantly <br> hydrogen <br> or <br> main sequence stars are rich in hydrogen $\checkmark$ | $\mathbf{2}$ |  |
| :--- | :--- | :--- | :--- | :--- | :--- |
| 16. | $\mathbf{a}$ | $\mathbf{i i}$ | peak wavelength: $290 \pm 10$ «nm» $\checkmark$ <br> $T=\frac{2.9 \times 10^{-3}}{290 \times 10^{-9}}=« 10000 \pm 400 \mathrm{~K} » \checkmark$ | Substitution in equation must be <br> seen. <br> Allow ECF from MP1. | $\mathbf{2}$ |

(Question 16 continued)

| 16. | b | i | $35 \pm 5 L_{\text {s }}$ V |  | 1 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 16. | b | ii | $\begin{aligned} & \frac{L_{\mathrm{x}}}{L_{\mathrm{s}}}=\frac{R_{\mathrm{x}}^{2} \times \mathrm{T}_{\mathrm{x}}^{4}}{R_{\mathrm{s}}^{2} \times \mathrm{T}_{\mathrm{s}}^{4}} \\ & \text { OR } \\ & R_{\mathrm{x}}=\sqrt{\frac{L_{\mathrm{x}} \mathrm{~T}_{\mathrm{s}}^{4}}{L_{\mathrm{s}} \mathrm{~T}_{\mathrm{x}}^{4}}} \times R_{\mathrm{s}} \checkmark \\ & R_{\mathrm{x}}=\sqrt{\frac{35 \times 6000^{4}}{10000^{4}}} \times R_{\mathrm{s}} \text { (mark for correct substitution) } \\ & R_{\mathrm{x}}=2.1 R_{\mathrm{s}} \end{aligned}$ | Allow ECF from (b)(i). <br> Accept values in the range: 2.0 to $2.3 R_{s}$. <br> Allow $T_{s}$ in the range: 5500 K to 6500 K . | 3 |
| 16. | b | iii | $\begin{aligned} & M_{\mathrm{x}}=(35)^{\frac{1}{3.5}} M_{\mathrm{s}} \checkmark \\ & M_{\mathrm{x}}=2.8 M_{\mathrm{s}} \checkmark \end{aligned}$ | Allow ECF from (b)(i). <br> Do not accept $M_{X}=(35)^{\frac{1}{3.5}}$ for first marking point. <br> Accept values in the range:2.6 to $2.9 M_{s}$. | 2 |
| 16. | C |  | the star «core» collapses until the «inward and outward» forces / pressures are balanced $\sqrt{ }$ <br> the outward force / pressure is due to electron degeneracy pressure «not radiation pressure» $\checkmark$ |  | 2 |


| 17. | a | experiments and collecting data are extremely costly $\checkmark$ data from many projects around the world can be collated $\checkmark$ | OWTTE | 1 max |
| :---: | :---: | :---: | :---: | :---: |
| 17. | b | $\begin{aligned} & v=« z c=0.19 \times 3 \times 10^{8}=» 5.7 \times 10^{7}<\mathrm{ms}^{-1} » \checkmark \\ & d=« \frac{v}{H_{0}}=\frac{5.7 \times 10^{4}}{70} »=810 \mathrm{Mpc} \text { OR } 8.1 \times 10^{8} \mathrm{pc} v \end{aligned}$ | Correct units must be present for MP2 to be awarded. <br> Award [2] for BCA. | 2 |
| 17. | c | ALTERNATIVE 1 $\frac{R_{\text {now }}}{R_{\text {then }}}=1+z=1.19 \mathrm{~V}$ <br> so (assuming constant expansion rate) $\frac{t_{\text {now }}}{t}=1.19 \checkmark$ $t=\frac{14}{1.19}=11.7 \mathrm{By}=12$ «By (billion years)» $\checkmark$ <br> ALTERNATIVE 2 <br> light has travelled a distance: $\left(810 \times 10^{6} \times 3.26=\right) 2.6 \times 10^{9} \mathrm{ly}$ V so light was emitted: 2.6 billion years ago $\checkmark$ so the universe was 11.4 billion years old $\checkmark$ | MP1 can be awarded if MP2 clearly seen. <br> Accept $2.5 \times 10^{25} \mathrm{~m}$ for MP1. <br> MP1 can be awarded if MP2 clearly seen. | 3 |


| 18. | a |  | a white dwarf accretes mass «from a binary partner» $\checkmark$ when the mass becomes more than the Chandrasekhar limit $\left(1.4 \mathrm{M}_{\mathrm{s}}\right)$ «then a supernova explosion takes place» $\checkmark$ |  | 2 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 18. | b | i | $\begin{aligned} & d=\sqrt{\frac{L}{4 \pi b}}=\sqrt{\frac{5 \times 10^{5} \times 3.8 \times 10^{26}}{4 \pi \times 1.6 \times 10^{-6}}} \\ & d=3.07 \times 10^{18} « \mathrm{~m} » \end{aligned}$ | At least 3 sig fig required for MP2. | 2 |
| 18. | b | ii | type la supernova can be used as standard candles $\checkmark$ there is no dust absorbing light between Earth and supernova $\checkmark$ their supernova is a typical type la $\checkmark$ |  | 1 max |


| 19. | a | $\frac{m v^{2}}{r}=\frac{G M m}{r^{2}}$ and correct rearranging $\checkmark$ | 1 |
| :---: | :---: | :---: | :---: |
| 19. | b | linear /rising until $\mathrm{R}_{0} \checkmark$ then «almost» constant $\checkmark$ | 2 |
| 19. | c | for $v$ to stay constant for $r$ greater than $R_{0}, M$ has to be proportional to $r \checkmark$ but this contradicts the information from the $M$-r graph $\checkmark$ OR <br> if $M$ is constant for $r$ greater than $R_{0}$, then we would expect $v \propto r^{\frac{-1}{2}} \checkmark$ but this contradicts the information from the $v$-r graph $\checkmark$ | 2 max |

